

60-GHz Millimeter-Wave Channel Measurements and Modeling for Indoor Office Environments

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Abstract—The millimeter-wave (mmWave) band will be used for the fifth-generation communication systems. In this paper, 60-GHz mmWave channel measurements and modeling are carried out for indoor office environments. The rotated directional antenna-based method and uniform virtual array-based method are adopted and compared to investigate the 60-GHz channel in a 3-D space, simultaneously covering azimuth and elevation domains. The multipath component parameters including power, delay, azimuth, and elevation angles are estimated with the space-alternating generalized expectation-maximization estimation algorithm, and then processed with the K-means clustering algorithm. An extended Saleh-Valenzuela model with both delay and angular cluster features is used to characterize the measured channel, and the intercluster and intracluster parameters are extracted. We find that the azimuth departure angles are diverse and highly related to the antenna position and measurement environment, while the elevation departure angles are more related to the antenna height difference and confined in a

relatively small direction range. The azimuth angle spread is much larger than the elevation angle spread either in global level or in cluster level. The results agree with the studies in the literature and channel models in IEEE standards.

Index Terms—60-GHz millimeter-wave (mmWave) channels, channel measurements, parameter estimation, Saleh-Valenzuela (S-V) channel model, space-alternating generalized expectation-maximization (SAGE) algorithm.

I. INTRODUCTION

WIRELESS communication networks are facing new challenges as Internet usages are proliferating. Future networks have to support high capacity, high mobility [1], and massive connectivity with increasingly diverse set of services, applications, and users [2], such as machine-to-machine and Internet of things, to enable communications anywhere, any-time, and by anything. Also, the flexible and efficient use of all spectrums and resources are required to accommodate widely different network deployment scenarios. These requirements have been envisioned for the research and development of the fifth-generation (5G) wireless communication systems. To address these challenges, there has been growing interest in moving up frequency into millimeter-wave (mmWave) bands where enormous bandwidths are available. MmWave technologies can achieve much higher data rate by using simple air interfaces due to the large bandwidth. In addition, by virtue of the vast available and unlicensed bandwidth, mmWave systems can satisfy higher system capacity requirement [3]. The mmWave band has drawn increasing attention despite earlier hiccups associated with the cost of producing commercial products [4]. With advances in mmWave chipset manufacturing led by, among others, Intel and IBM, this band has been opened up by various regulatory authorities in the process of developing the first 60-GHz mmWave standards and specifications such as IEEE 802.15.3c [5] for wireless personal area networks and IEEE 802.11ad [6] for multigigabit wireless local area networks.

The propagation characteristics at mmWave bands are quite different from those at frequency bands below 6 GHz. Because of small wavelengths at mmWave bands, diffraction becomes the weakest and least reliable propagation mechanism for mmWave communication systems. Many small objects in the environment can be considered as scatterers, thus exhibiting a highly reflective and scattering nature to create alternative viable links, except for the direct link [7]. Also, atmospheric attenuation caused by air and water molecules creates additional path loss (PL) at mmWave frequencies, especially at

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60 GHz. In addition, with identical antenna gains and transmit powers, the power received at mmWave bands is smaller than that at lower frequencies due to the smaller antenna aperture. Due to these propagation characteristics, highly directional antennas, large antenna array, and adaptive beamforming techniques may be used to establish reliable communication links in mmWave wireless communication systems [8].

Recently, there have been extensive research on mmWave outdoor cellular communications. In [9]–[11], the measurements were conducted at 28, 38, and 73 GHz in urban outdoor environments. These studies mainly focused on large-scale properties of wireless channels. Results show that by using directional or adaptive beamforming antennas, mmWave signals are potentially viable at distances of 100–200 m, even in nonline-of-sight (NLOS) scenarios. Besides, mmWave systems can provide at least an order of magnitude improvement in capacity over long-term evolution systems, at least for outdoor coverage.

Current cellular communication systems normally use an outdoor base station (BS) in the middle of a cell to communicate with mobile users, no matter whether they stay indoors or outdoors. For indoor users communicating with an outdoor BS, the signals have to go through building walls, which cause very high penetration loss that significantly degrades the data rate, spectral efficiency, and energy efficiency of wireless transmissions. At mmWave frequencies, this situation will be even worse. Therefore, outdoor and indoor scenarios will be probably separated in 5G cellular architecture [12].

Extensive channel measurement campaigns at mmWave frequencies have been conducted in indoor environments, especially at 60 GHz [13]–[16]. Large-scale properties like PL and delay dispersion at mmWave frequencies in indoor channels have been studied thoroughly [15], [17], [18]. Studies show that the PL exponent n varies from 0.40 to 2.10 and from 1.97 to 5.40 for line-of-sight (LOS) and NLOS scenarios, respectively, in various 60-GHz indoor environments. Specifically, in a typical office environment, n takes values in the ranges from 1.16 to 2.17 and from 2.83 to 4.00 for LOS and NLOS scenarios, respectively [19]. Delay dispersion is affected by the dimensions of the room, the wall reflection coefficient, and antenna polarization and beamwidth [13], [20]. With respect to channel modeling for 5G mmWave indoor communication systems, the characteristics of both azimuth and elevation angles in 3-D space, large antenna array, super-high resolution in delay domain, and small-scale fading characteristics should be considered.

Angular characteristics of multipath propagation are critical for mmWave communication systems that exploit highly directional steering antennas and adaptive array. In general, there are two kinds of measurement methods for indoor mmWave channel angular characterization. One is the rotated directional antenna (RDA)-based method and the other is the uniform virtual array (UVA)-based method. Both methods are only applied to quasi-static environment due to the long time required for the antenna to scan directions or shift positions in the 3-D space. In the RDA-based method, a highly directional antenna is used with a vector network analyzer (VNA) or channel sounder. The antenna is placed

on the positioner and scanned in the azimuth and elevation angular domains with a small rotation angle step during the measurements. Because of the narrow beamwidth of the antenna, the antenna pointing angle is taken as the angle of arrival (AoA) or angle of departure (AoD) of multipath components (MPCs) [21]. In the UVA-based method, the RDA is replaced by an omnidirectional antenna which shifts position in the 3-D space with a step less than a half wavelength [22]. For the estimation of UVA measurement channel MPC parameters, high-resolution parameter estimation algorithms like subspace-based multiple signal classification and estimation of signal parameters via rotational invariance technique algorithms, and maximum-likelihood (ML)-based space-alternating generalized expectation-maximization (SAGE) and Richter's ML parameter estimation framework algorithms are usually utilized to process the measurement data.

Two IEEE standards, 802.15.3c and 802.11ad, have been proposed for 60-GHz indoor wireless communications. IEEE 802.15.3c is a single-input multiple-output channel model and only characterizes the AoA in azimuth domain, while IEEE 802.11ad is a multiple-input multiple-output (MIMO) channel model and characterizes the double-directional angle properties. A detailed comparison of the channel models in two standards can be found in [23]. The channel models for both standards are based on the angular cluster extended Saleh–Valenzuela (S-V) model [24]. In this model, MPCs are clustered in both time and angular domains. The intercluster and intracluster characteristics are modeled statistically in [25] and [26]. The detailed S-V-based channel model parameterization at 60 GHz in various indoor environments can be found in [27]–[30]. Because high-gain directional antennas are supposed to be used at mmWave frequencies, further investigation of directional characterization is also required. Although the IEEE 802.15.3c and IEEE 802.11ad models provide certain directional information, additional measurements for verification in a variety of environments would be desirable. Also, current studies on directional characteristics mainly focus on azimuth domain, but elevation domain characteristics are also important due to the use of massive MIMO and large antenna array at mmWave frequencies. The contributions of this paper are as follows.

- 1) We investigate the 60-GHz indoor wireless channel in the 3-D space, which simultaneously covers azimuth and its coelevation domains using both the RDA-based method and UVA-based method. Detailed measurement setup and full comparison of the two methods are given. Good agreement is shown between the RDA measurement results and UVA estimation results in both LOS and NLOS scenarios. The estimated MPCs using the SAGE algorithm are presented.
- 2) We provide a 60-GHz channel model and its parameterization for office environment based on the angular extended S-V model and temporal and spatial clustering properties.
- 3) We model the channel in both intercluster and intracluster levels, especially in angular domains.

The remainder of this paper is organized as follows. Section II describes the measurement environment, mea-

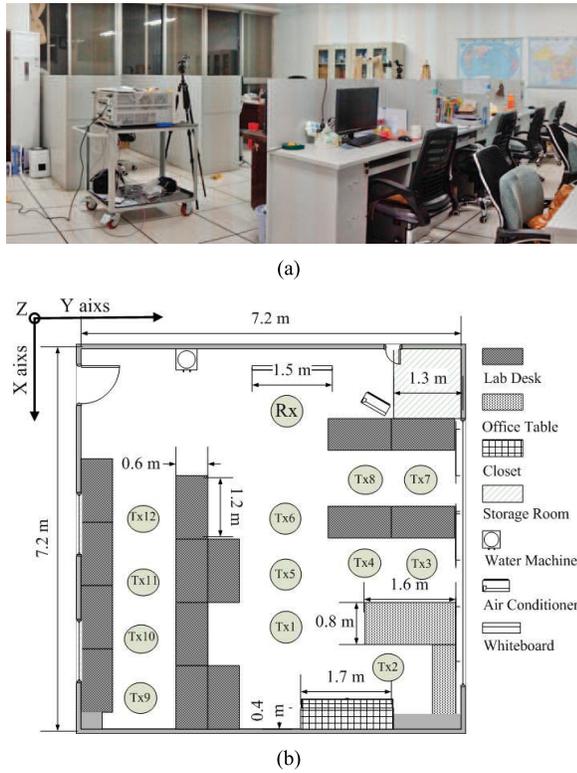


Fig. 1. (a) Photo of the office environment. (b) Layout of the office.

surement system setup, channel measurement methods, measurement methods' comparison, and SAGE algorithm. In Section III, the measured and estimated channel results for RDA- and UVA-based methods are presented and fully compared. The channel modeling and evaluation from the measurement results are described in Section IV. Finally, conclusions are drawn in Section V.

II. 3-D INDOOR MMWAVE CHANNEL MEASUREMENTS

A. Measurement Environment

The indoor measurements are conducted in an office environment as shown in Fig. 1(a), while the layout is shown in Fig. 1(b). The size of the room is about $7.2 \times 7.2 \times 3$ m³. The office is furnished with multiple chairs, desks, and a table. The desks are about 0.75-m height at desktop level but with an additional vertical clapboard having about 0.45-m height. Thus, the total height is about 1.2 m. In addition, the desks and table are equipped with several computers and electronic devices. Other large objects include a closet, an air conditioner, a water machine, and a whiteboard. The walls, floor, and ceiling are made of concrete. Parts of the floor and ceiling are made of antistatic-electricity board. There are several windows on both sides of the wall. The locations of the 12 transmitter (Tx) and the receiver (Rx) in the UVA measurement are also labeled in Fig. 1(b). In the RDA measurement, the Rx position is the same with the UVA measurement while the Tx is located at Tx1 and Tx3.

B. Measurement System Setup

The measurement system setup is shown in Fig. 2. A PC is used to control the antenna positioner through a universal

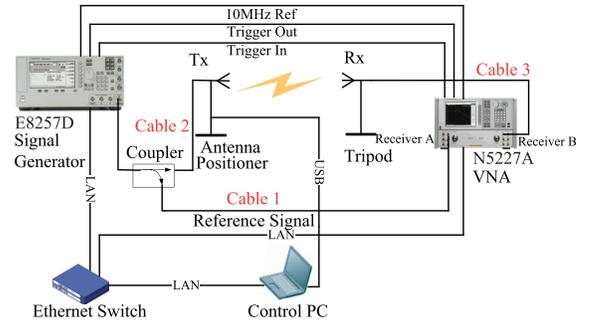


Fig. 2. Measurement system setup.

serial bus port. The PC also controls the Keysight E8257D analog signal generator and Keysight N5227A VNA through local area network ports. The analog signal generator and VNA are connected by 10-m-long trigger in/out and 10-MHz time base cables, thus achieving precise synchronization. With the signal generator, the measurable distance can be extended, and the measurement system does not need power amplifier, which may introduce noises into the measured channel data. The signal generator is used as the source whose output signal is split into two parts by a directional coupler. One part is fed to the Tx antenna as a channel sounding signal, and the other part is fed to the VNA as the reference signal. The VNA is used as the Rx which is connected with the Rx antenna by a coaxial cable. The customized high-precision antenna positioner has 5 degrees of freedom in the $X/Y/Z$ /azimuth/elevation axes. Its scan ranges are -180° to 180° in azimuth, 0° – 180° in elevation, ± 300 mm in X -axis, 0 – 600 mm in Y -axis, and ± 250 mm in Z -axis, as shown in Fig. 3. The antenna positioner is placed on the Tx side. With the antenna positioner, Tx antenna can scan in azimuth and elevation 3-D angular domains for the RDA measurements and scan in $X/Y/Z$ axes to form a 3-D virtual antenna array for the UVA measurements. At the Rx side, a tripod is used to support a fixed horn antenna. We define that the azimuth angle ϕ along X -axis is 0° . It increases when rotating counterclockwise and decreases when rotating clockwise. Therefore, $\phi \in [-180^\circ, 180^\circ]$ is held. The elevation angle θ is defined to be 0° along Z -axis and it increases when rotating downward as shown in Fig. 3. So, $\theta \in [0^\circ, 180^\circ]$ is held.

C. Channel Measurement Methods

1) *RDA Method*: The RDA-based measurement approach is to steer the highly directional antenna in azimuth and elevation domains in the measurement as shown in Fig. 4. This method has been used in [9], [10], and [21] for outdoor urban measurements, but the elevation pointing angles were very limited, for example, just pointed to 0° and $\pm 10^\circ$. The narrow beamwidth directional antenna divides the space into multiple sectors. In the measurement, both Tx and Rx antennas are 25-dBi standard gain horn antennas with E- and H-plane 3-dB beamwidth of 10° at 60 GHz. The Tx antenna is placed on the antenna positioner which is rotated from -180° to 180° in azimuth angles and from 30° to 150° in elevation angles with the step of 5° in one snapshot of the channel measurements. Note that the intention of RDA method is to

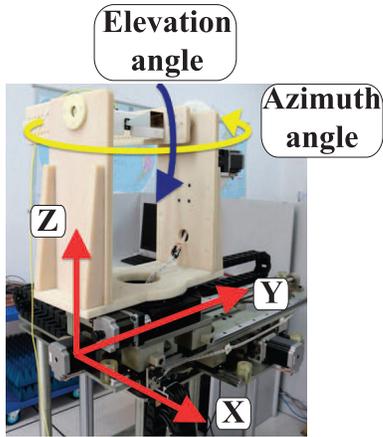


Fig. 3. Antenna positioner with five axes ($X/Y/Z$ /azimuth/elevation).

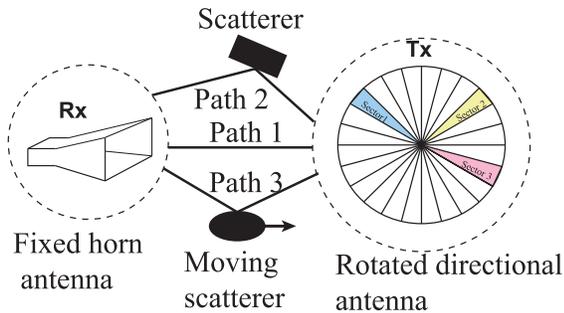


Fig. 4. RDA-based measurement method.

achieve the omnidirectional coverage with high gain, and the antenna beamwidth and rotation step will have impact on the measurements. Actually, the same MPCs will appear across multiple adjacent angular bins, which will make troubles when discussing power angle spectrum (PAS). However, when we deal with the power delay profile (PDP), superposition of PDPs in every direction is equivalent to create a nearly flat antenna power response over all spatial directions. As shown in [31], the same 10° beamwidth horn antenna was rotated with step size of 10° and the virtual antenna array was formulated. It demonstrated that the array magnitude fluctuates in azimuth are within 0.25 dB, and such a variation is considered small and negligible. When the rotation step is smaller, the fluctuates will be smaller, and the absolute gain will go higher. The gain will be deducted from the final PDP. To have a fair comparison of the propagation characteristics for the twelve links and see the coverage range of the Rx antenna, the Rx antenna is placed on a tripod with a fixed location and pointing angle along X -axis during the whole channel measurement. The heights of Tx and Rx antennas are both 1.6 m. The Tx antenna is located at Tx1 and Tx3 which correspond to LOS and NLOS scenarios, respectively. During the measurement, the environment is kept static. The measured frequency range is from 59 to 61 GHz using 401 frequency points. The delay resolution is inverse to the measurement bandwidth. In our measurement, the delay resolution is 0.5 ns, which means that two paths with delay difference below 0.5 ns are not

TABLE I
MEASUREMENT SETTINGS FOR THE RDA-BASED MEASUREMENT

| Center frequency | Bandwidth | Sampling points | Output power |
|------------------|---|-------------------------------|---------------------|
| 60 GHz | 2 GHz | 401 | +13 dBm |
| IF bandwidth | Tx locations and scan points in azimuth and elevation | Rx antenna gain and beamwidth | Rotation step angle |
| 1 kHz | Tx1 and Tx3, 72×25 | 25 dBi, 10° | 5° |

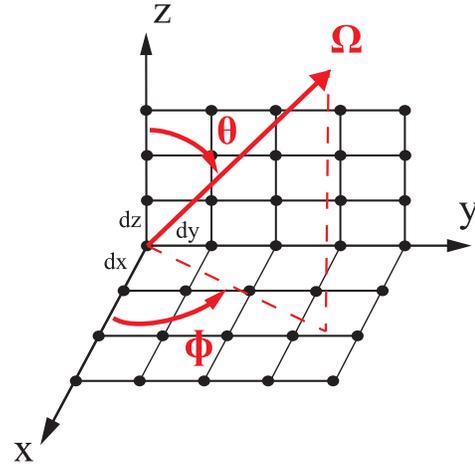


Fig. 5. UVA configuration and direction characterization.

resolvable. This is the delay resolution for Fourier transform. The output power of the signal generator is 13 dBm and the intermediate frequency (IF) filter bandwidth is 1 kHz. Therefore, the obtained data set from the measurement is a $25 \times 72 \times 401$ matrix. The key settings for this measurement are shown in Table I.

2) *UVA Method*: The system setup for the UVA-based channel measurement is similar to the RDA method. The Tx antenna is replaced by a hollow cube UVA formed of a monopole omnidirectional antenna. The configuration of the UVA is illustrated in Fig. 5. This method can be found in [22] and [23], but the virtual antenna array size was limited to $4 \times 4 \times 4$ in [22] and 7×7 in [23]. Because of the symmetry of the steering vector, the horizontal virtual rectangular array cannot distinguish impinging waves from above and below, and the vertical virtual rectangular array cannot distinguish impinging waves from front and back. Therefore, the UVA adopted in the measurement is formed of six surfaces of a $15 \times 15 \times 15$ cube. Each surface is either a horizontal or vertical virtual rectangular array consisting of 15×15 antenna elements. With this large 3-D virtual antenna array, the precise and high-resolution angle parameters can be estimated using high-resolution parameter estimation algorithms. The spacing between adjacent positions in each surface is 2.5 mm, i.e., a half wavelength at 60 GHz. More comprehensive channel measurement campaigns have been conducted for this method. The Rx antenna is the same 25-dBi horn antenna. The heights of Tx and Rx antennas are both 1.6 m. Twelve Tx locations are selected inside the room as shown in Fig. 1(b), which are positions of the user equipment and boresight LOS links with different Tx–Rx distances. The measured frequency range is

TABLE II
MEASUREMENT SETTINGS FOR THE UVA-BASED MEASUREMENT

| Center frequency 60 GHz | Bandwidth 2 GHz | Sampling points 401 | Output power +13 dBm |
|----------------------------|--|--|----------------------------|
| IF bandwidth 1 kHz | Tx locations and shift points in cube array Tx1-Tx12, 15 × 15 × 6 | Rx antenna gain and beamwidth 25 dBi, 10° | Position spacing 2.5 mm |

from 59 to 61 GHz and 401 frequency samples have been collected. The obtained data set from the measurement is a $15 \times 15 \times 6 \times 401$ matrix for each Tx position. During the measurement, the environment is kept static. The key settings for this UVA-based measurement are shown in Table II.

D. Measurement Methods' Comparison

The RDA-based method heavily relies on the accuracy of the measured radiation pattern of the directional antenna, while the rotation step is determined by the beamwidth of the radiation pattern. It may require long measurement time if very directional antenna is used. The highly directional antenna may not be rotated with its feeding port fixed exactly during the whole measurement. Thus, the antenna radiation pattern observed when the antenna points toward a certain direction may be different from the original radiation pattern with axis shifted. Furthermore, the cable connected to the antenna may introduce additional nonuniform phase drifts when the antenna is rotated, which is hard to calibrate [31]. But this method can provide large dynamic range of the received signal power because of high gain of the directional antenna, which is likely to be used for practical mmWave wireless communication systems with beamforming.

For the UVA-based method, due to the use of the monopole antenna, the accuracy of the radiation pattern is not so crucial and the implementation of this method is relatively easy. Though this method provides smaller dynamic range in comparison with the RDA-based method for a single PDP measurement, the averaging over the antenna array can provide an antenna array gain. The advantage of this method is that through high-resolution parameter estimation algorithms, the estimated MPC parameters can achieve high precision and resolution in angle domains, for example, 1° in azimuth and elevation angles, by using a large 3-D virtual antenna array.

E. SAGE Algorithm

The SAGE algorithm has been studied extensively to jointly estimate high-resolution parameters of MPCs including the relative delay, AoA, AoD, Doppler frequency, and complex amplitude of impinging waves in mobile radio environments [32]. In a multiple-input single-output (MISO) scenario, the Tx is equipped with an antenna array consisting of M elements located at $r_1, \dots, r_M \in \mathbb{R}^2$ with respect to an arbitrary reference point. The underlying channel gain is assumed to consist of a finite number L of specular plane waves impinging at the Rx location. The received signal vector is [32]

$$Y(t) = \sum_{l=1}^L s(t; \theta_l) + \sqrt{\frac{N_0}{2}} N(t) \quad (1)$$

where $N(t)$ is a standard M -D vector-valued complex white Gaussian noise with spectral height of N_0 . Here, $s(t; \theta_l)$ is the signal contributed by the l th path at the output of the Rx array, i.e., an MPC, and can be expressed in vector notation [32]

$$\begin{aligned} s(t; \theta_l) &\triangleq [s_1(t; \theta_l), \dots, s_M(t; \theta_l)]^T \\ &= c(\Omega_l) \alpha_l \exp\{j2\pi \nu_l t\} u(t - \tau_l) \end{aligned} \quad (2)$$

where $\theta_l \triangleq [\tau_l, \Omega_l, \nu_l, \alpha_l]$ is the vector containing the parameters of the l th MPC, i.e., relative delay τ_l , AoD Ω_l , Doppler frequency ν_l , and complex amplitude α_l . Because the measured channel is static, Doppler frequency is not considered in the proposed SAGE algorithm. In (2), $u(t - \tau_l)$ is the delayed version of Tx signal. The M -D vector-valued function $c(\Omega) \triangleq [c_1(\Omega), \dots, c_M(\Omega)]^T$ is the steering vector of the array and its components are given by [32]

$$c_m(\Omega) \triangleq f_m(\Omega) \exp\{j2\pi \nu_l t \lambda^{-1} \langle e(\Omega), r_m \rangle\}, \quad m = 1, \dots, M \quad (3)$$

with λ , $e(\Omega)$ and $f_m(\Omega)$ denoting the wavelength, the unit vector in \mathbb{R}^2 pointing toward the direction determined by Ω , and the complex electric field pattern of the m th antenna element, respectively. The SAGE algorithm is a variant of the expectation-maximization algorithm which consists of two steps: E-step and M-step. More details and implementation of the SAGE algorithm can be found in [32] and [33]. The focus of this section is the generation of steering vector of the array for different measurement methods.

For the UVA-based method

$$\langle e(\Omega), r_m \rangle = [x_m - d_x, y_m - d_y, z_m - d_z] e(\Omega) \quad (4)$$

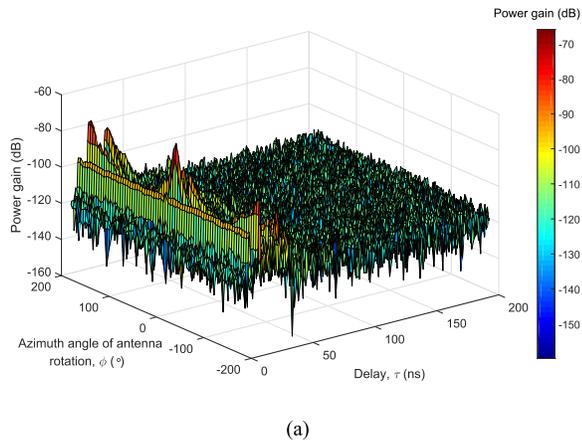
where

$$e(\Omega) = [\cos(\phi) \sin(\theta), \sin(\phi) \sin(\theta), \cos(\theta)]^T. \quad (5)$$

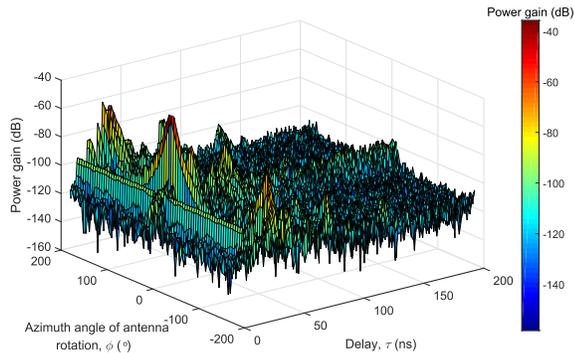
The direction Ω is defined as a unit vector with initial point anchored at the origin of coordinate system and it is uniquely determined by its spherical coordinates $(\phi, \theta) \in [-180^\circ, 180^\circ] \times [0^\circ, 180^\circ]$ as shown in Fig. 5. The angles ϕ and θ are referred as the azimuth and the coelevation of Ω , respectively, while x_m , y_m , and z_m are the position coordinates of the m th antenna element. The spacing between adjacent position in each axis is half wavelength at 60 GHz, i.e., $d_x = d_y = d_z = 2.5$ mm.

For the RDA-based method, because the antenna is rotated coaxially, the spacing between different positions is assumed to be zero. Therefore, the steering vector for this method is equal to the complex electric field pattern of the m th antenna element, i.e., $c_m(\Omega) = f_m(\Omega)$.

In the literature, the RDA-based measurement data is often processed with simple peak search algorithm and the rotated angle is assumed to be the path AoA or AoD. This is heavily affected by the antenna characteristics. High-resolution algorithms like SAGE may be affected by phase inaccuracy, and the acquire of 3-D antenna radiation pattern in 60-GHz frequency band is also hard and time consuming. Due to the lack of accurate antenna radiation pattern, the RDA measurement data are not processed with SAGE algorithm in this paper.



(a)



(b)

Fig. 6. Measured delay-AS function at Tx1. (a) Elevation angle = 30° . (b) Elevation angle = 90° .

III. CHANNEL MEASUREMENT RESULTS

A. RDA-Based Measurement Results

From the measurement, 1800 instantaneous directional PDPs are obtained which correspond to different Tx antenna pointing angles in azimuth and elevation. The measured delay-azimuth spread (AS) function results for 25-dBi Tx antenna at Tx1 with different elevation angles for the 60-GHz indoor channel using rotated horn antenna are shown in Fig. 6. The LOS component and MPCs are clearly identified from the measured delay-AS function. Some of the PDPs are shown in Fig. 7. The dynamic range of the measurement system is about 60 dB. A LOS component can be clearly observed when two antennas are pointing to each other with the relative received power of up to -40 dB. When antennas are misaligned, the power is reduced gradually.

To the best of our knowledge, the comparison of the RDA measurement results and UVA estimation results has never been done in the literature. To do the comparison, the RDA measurement results are compared with the UVA measurement results at Tx1 and Tx3. For LOS scenario at Tx1, the average PDPs obtained by the RDA measurement, UVA measurement, and UVA estimation are shown in Fig. 8. For the RDA measurement, the 25-dBi Tx antenna gain has been subtracted. The two average PDPs show good agreement with each other. The LOS component arrives with relative received power about -60 dB and delay of 13.5 ns, which corresponds to 4-m

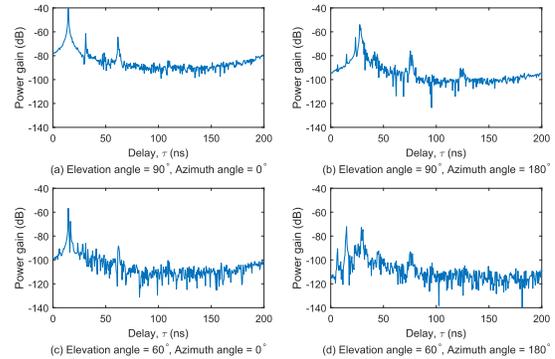


Fig. 7. Different directional PDPs at Tx1. (a) Elevation angle = 90° , azimuth angle = 0° . (b) Elevation angle = 90° , azimuth angle = 180° . (c) Elevation angle = 60° and azimuth angle = 0° . (d) Elevation angle = 60° , azimuth angle = 180° .

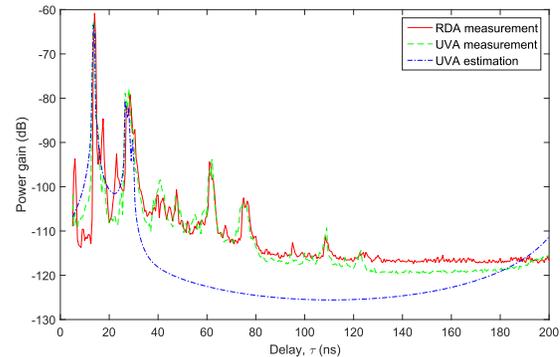


Fig. 8. Average PDP for RDA- and UVA-based measurements at Tx1.

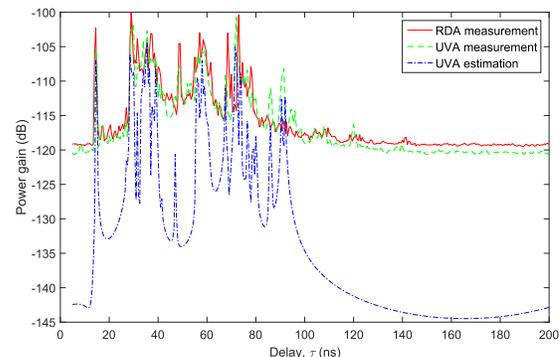
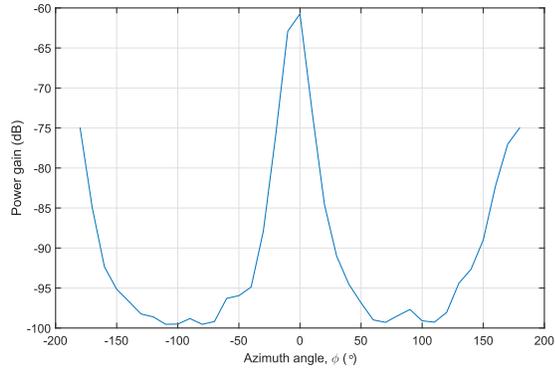
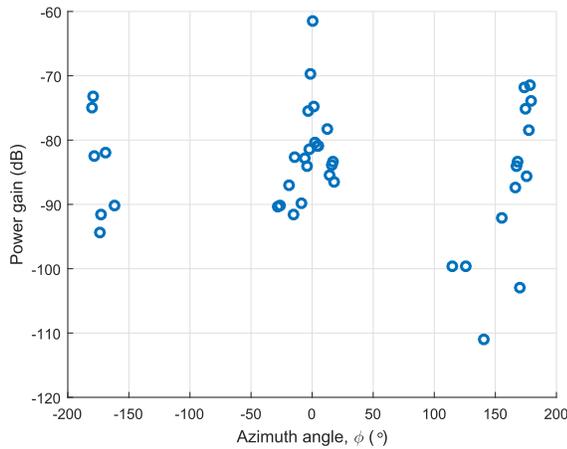


Fig. 9. Average PDP for RDA- and UVA-based measurements at Tx3.

separation. Another first-order reflection component at 26.5 ns in delay can also be seen with the power degradation of 20 dB compared with the LOS component. The second-order reflection at about 60 ns in delay has the power degradation of 34 dB with respect to the LOS component. Due to its weakness, this path is not estimated by the SAGE algorithm of 30-dB threshold. Later we will show that this is also due to the fact that the strong path has interference on the estimation of weak path in LOS scenario, while in NLOS scenario, all paths can be accurately estimated even with relatively weak powers, as shown in Fig. 9 for Tx3. The power azimuth profile (PAP) comparison at Tx1 is shown in Fig. 10. The PAP is obtained by summing linear power over all elevation angles for each azimuth angle using the RDA measured data,



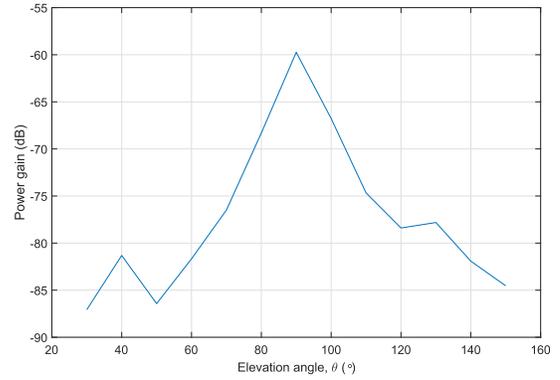
(a)



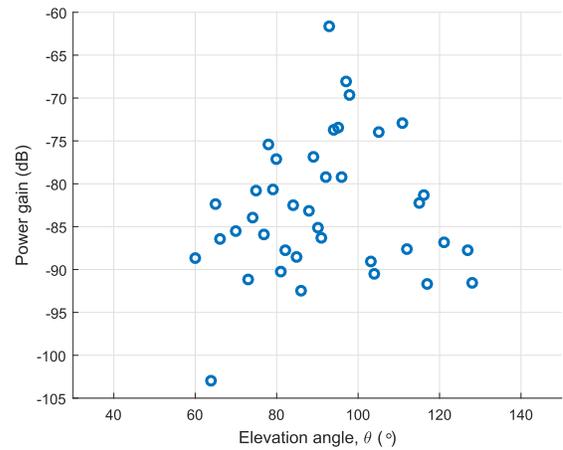
(b)

Fig. 10. PAP comparison at Tx1. (a) RDA measurement. (b) UVA estimation.

while it is obtained by combining the SAGE estimated MPCs with the same azimuth angle but differ in delays and elevation angles. Although objects placed near the Tx may have higher order reflections and diffractions, the relative received power is relatively weak and these MPCs are not dug out by the SAGE algorithm. The power elevation profile (PEP) comparison at Tx1 is shown in Fig. 11. The PEP is obtained by summing linear power over all azimuth angles for each elevation angle, while it is obtained by combining the SAGE estimated MPCs with the same elevation angle but differ in delays and azimuth angles. We can see that the PAP and PEP obtained by the two methods match well. The LOS component departs with 0° in azimuth angle and 90° in elevation angle. The first-order reflection component departs with $\pm 180^\circ$ in azimuth angle while confined in a narrow direction range near 90° in elevation angle. However, some little differences can be seen. The UVA estimation results show relatively small elevation angle range about 60° – 130° compared with the RDA measurement which shows angle range about 40° – 140° with relatively strong power. The normalized PAS comparison at Tx1 is shown in Fig. 12. The PAS is obtained from the measured 1800 instantaneous directional PDPs by summing over delay domain for RDA measured data and the SAGE estimates from the UVA measured data, respectively. For RDA method, the MPC AoDs are using the antenna pointing angles on the positioner from -180° to 180° in azimuth and from



(a)



(b)

Fig. 11. PEP comparison at Tx1. (a) RDA measurement. (b) UVA estimation.

30° to 150° in elevation with a step of 5° . Except the LOS component at (azimuth angle = 0° , elevation angle = 90°), a major scatterer MPC can be observed at (azimuth angle = $\pm 180^\circ$, elevation angle = 90°). Overall, the results obtained by the two methods show good agreement with each other in terms of the average PDP, PAP, PEP, and PAS, which verifies the validation of the two measurement methods.

B. UVA-Based Measurement Results

In SAGE algorithm, the number of MPCs is usually predefined which is large enough to capture all the significant paths. In the literature, this large number ranges from 50 to 400. In this paper, we have changed the estimated number of paths and find out that 100 paths are appropriate to capture most of the signal energy while within an acceptable computation complexity.

Twelve Tx locations are selected in this set of measurements including LOS scenarios and NLOS scenarios. The power, AoD in azimuth and elevation, and delay of MPCs are all estimated. The K-means algorithm [34] is adopted for clustering based on delay, AoD in azimuth and elevation domains of MPCs estimated from the SAGE algorithm. A KPowerMeans algorithm was introduced in [35] based on the K-means algorithm, and considered power weight to calculate the cluster centroid and used MPC distance as a metric to combine

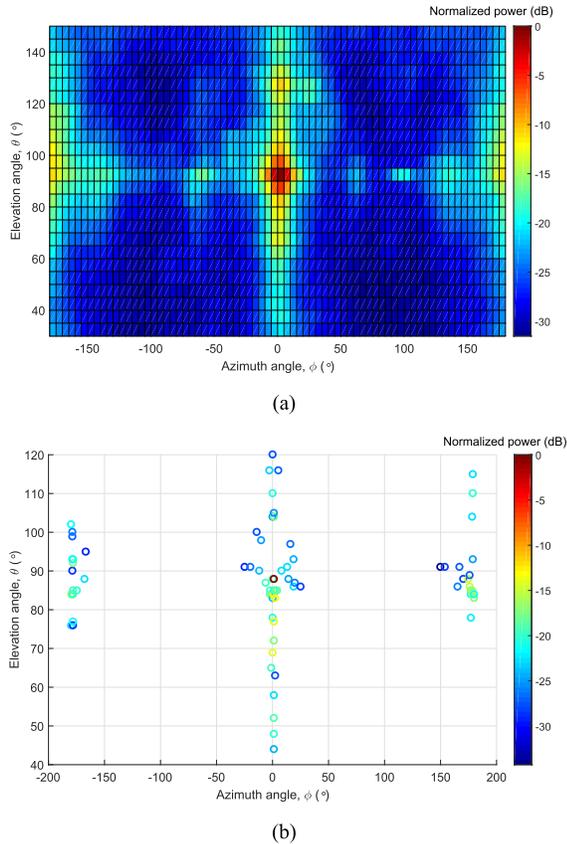


Fig. 12. PAS comparison at Tx1. (a) RDA measurement. (b) UVA estimation.

normalized delays and angles with different units. However, the delay scale parameter is unknown and related to the real measurement data set and can highly affect the clustering results. In addition, in the extended S-V model, the powers of clusters and rays in each cluster have a decay rate, implying that powers may change greatly within a cluster. Thus, we use the traditional K-means algorithm to do the clustering instead.

There are some scientific criteria for determining the number of clusters, among which the Kims–Park (KP) index, Calinski–Harabasz index, Xie–Beni index, and PBM index are most popularly used [36], [37]. Though there are so many scientific criteria, several studies on performance comparison showed that there is no single index which outperforms all the others [37]. Besides, the optimal number may be different according to the data type in LOS and NLOS scenarios. Improvements may be achieved by combining several indices, but the weight of each index is an open question. In this paper, we chose the KP index to determine the optimal cluster number. A cluster is defined as a group of MPCs showing similar delay and AoD. Clusters should not overlap. For each identified cluster i , a set of parameters $C_i = (T_i, \Phi_i, \Theta_i)$ is used to approximate the center of the cluster, where T_i , Φ_i , and Θ_i are delay, AoD in azimuth domain, and AoD in elevation domain, respectively. The center of a cluster is defined as the mean value over each domain. As shown in Fig. 13 for Tx1 position with Tx–Rx separation of 4 m, six clusters are observed, centered at approximately $C_1 = (14 \text{ ns}, 178^\circ, 85^\circ)$, $C_2 = (13.5 \text{ ns}, 0^\circ, 109^\circ)$, $C_3 = (13.5 \text{ ns}, 1^\circ, 87^\circ)$, $C_4 = (13.5 \text{ ns}, 1^\circ, 57^\circ)$, $C_5 = (13.5 \text{ ns}, -178^\circ, 84^\circ)$, and

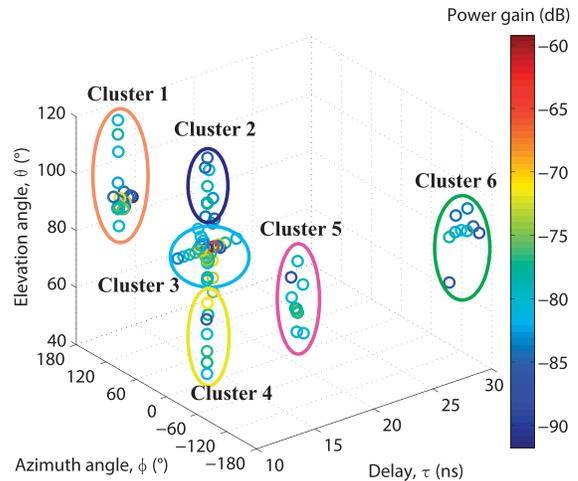


Fig. 13. MPCs and clustering results for UVA-based measurement at Tx1.



Fig. 14. Measurement setup for UVA-based measurement at Tx1.

$C_6 = (26.5 \text{ ns}, -178^\circ, 91^\circ)$. Cluster 3 is the LOS component, although it has wide angle spread in azimuth domain. From clusters 1, 2, 4, and 5 shown in Fig. 13, it can be observed that some MPCs have very different AoDs (sometimes larger than 100°) but quite close delays (less than 1 ns). Note that the antenna positioner has effects on the measured channel. This effect is hard to remove and may be modeled as stochastic scatterers. For example, the same situation can be observed in cluster 1, where the signal is reflected by the antenna support on the positioner as shown in Fig. 14. Cluster 6 is at approximately 26.5 ns which is the first-order reflection from the wall based on wave traveling distance and AoD information. Further investigation of the estimated MPCs shows that for the three boresight LOS scenarios, i.e., Tx1, Tx5, and Tx6, each position shows the two strong clusters departure with azimuth angle near 0° and $\pm 180^\circ$, but with different delays according to different Tx antenna positions.

Fig. 15 shows the estimated MPCs and their clustering for Tx11. Six clusters are identified and their centers are approximately at $C_1 = (33 \text{ ns}, -39^\circ, 91^\circ)$, $C_2 = (31 \text{ ns}, -164^\circ, 90^\circ)$, $C_3 = (58 \text{ ns}, 79^\circ, 98^\circ)$, $C_4 = (58 \text{ ns}, -8^\circ, 87^\circ)$, $C_5 = (58 \text{ ns}, -170^\circ, 98^\circ)$, and $C_6 = (78.5 \text{ ns}, -174^\circ, 91^\circ)$. The distance between the Tx and Rx is 4 m. According to delays of MPCs, there is no LOS component. This is because of the narrow beamwidth of horn antenna at the Rx. Fig. 16 shows the synthesized PDP from the SAGE results, which collects the total power of all rays in one delay bin. It can be

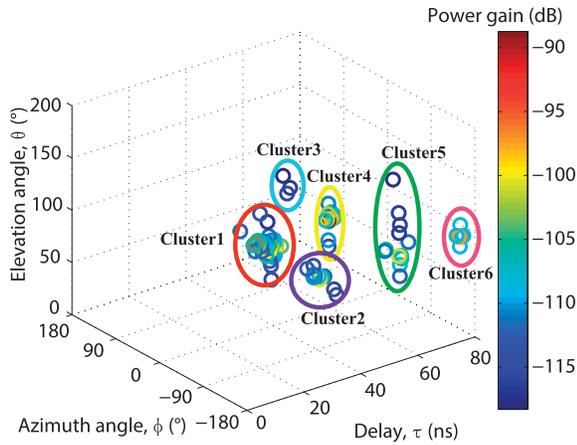


Fig. 15. MPCs and clustering results for UVA-based measurement at Tx11.

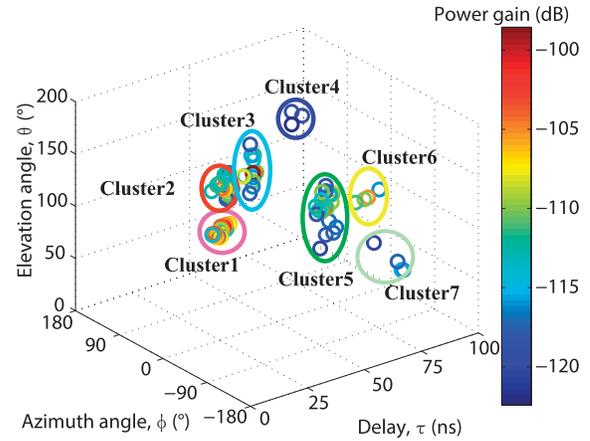


Fig. 17. MPCs and clustering results for UVA-based measurement at Tx3.

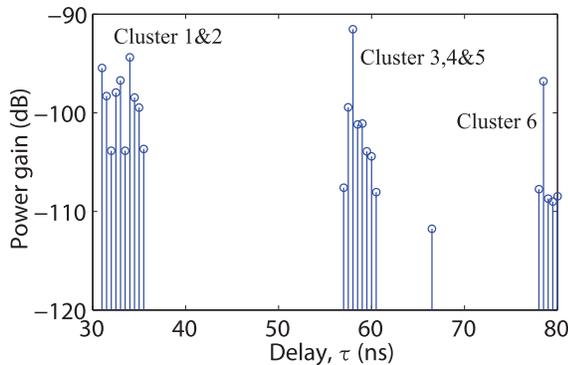


Fig. 16. Synthesized PDP for UVA-based measurement at Tx11.

observed that clusters 1 and 2 have very close delays, so as clusters 3–5. This is probably caused by scattering from the antenna positioner. For cluster 6, its delay is about 78.5 ns, which indicates multiple reflections or scattering. In terms of power, the maximum received single ray power is -88.7 dB in cluster 4. Even having multiple reflection/scattering, the peak single ray power in cluster 6 is up to -97 dB.

Fig. 17 shows the estimated MPCs and their clustering for Tx3, which is an NLOS scenario. Seven clusters are identified. Their centers are approximately at $C_1 = (33.5 \text{ ns}, 42^\circ, 89^\circ)$, $C_2 = (58 \text{ ns}, 158^\circ, 87^\circ)$, $C_3 = (72 \text{ ns}, 170^\circ, 88^\circ)$, $C_4 = (72 \text{ ns}, 57^\circ, 169^\circ)$, $C_5 = (72.5 \text{ ns}, 11^\circ, 89^\circ)$, $C_6 = (91.5 \text{ ns}, 8^\circ, 88^\circ)$, and $C_7 = (72 \text{ ns}, -174^\circ, 74^\circ)$. The Tx–Rx distance in this scenario is about 4.2 m. Because the Tx and Rx antennas are at the same height, most of elevation angles of MPCs are around 90° .

The SAGE estimated MPC parameters for all the 12 Tx positions show that azimuth AoD is diverse over -180° – 180° and largely related to Tx antenna position and the measurement environment, while elevation AoD is more related to the different heights of Tx and Rx antennas and confined in a relatively small direction range over 40° – 120° .

IV. 60-GHZ INDOOR CHANNEL MODELING

Due to the limited measurement data in RDA measurements, the 60-GHz indoor channel modeling is mainly based on

the measured data using the UVA method. Due to large bandwidths and therefore high delay resolution in mmWave 5G communication systems, delay and angular domain cluster characteristics have been observed in our measurement results, as well as in some other measurements in the literature. The original S-V model assumes that rays arrive in clusters in the delay domain, which can easily be extended to the angular domain. In this paper, we have therefore used the extended S-V model with both delay and angular cluster characteristics.

The channel impulse response for the channel model in MISO scenario without polarization characteristics is given by [25]

$$h(t, \phi_{Tx}, \theta_{Tx}) = \sum_i A^{(i)} C^{(i)}(t - T^{(i)}, \phi_{Tx} - \Phi_{Tx}^{(i)}, \theta_{Tx} - \Theta_{Tx}^{(i)}) \quad (6)$$

with

$$C^{(i)}(t, \phi_{Tx}, \theta_{Tx}) = \sum_k \alpha^{(i,k)} \delta(t - \tau^{(i,k)}) \cdot \delta(\phi_{Tx} - \phi_{Tx}^{(i,k)}) \delta(\theta_{Tx} - \theta_{Tx}^{(i,k)}) \quad (7)$$

where

ϕ_{Tx} and θ_{Tx} are azimuth and elevation angles at the Tx, respectively, $A^{(i)}$ and $C^{(i)}$ are the channel coefficient and channel gain for the i th cluster, respectively, $\delta(\cdot)$ is the Dirac delta function, $T^{(i)}$, $\Phi_{Tx}^{(i)}$, and $\Theta_{Tx}^{(i)}$ are time-angular coordinates of the i th cluster, $\alpha^{(i,k)}$ is the amplitude of the k th ray of the i th cluster, and $\tau^{(i,k)}$, $\phi_{Tx}^{(i,k)}$, and $\theta_{Tx}^{(i,k)}$ are relative time-angular coordinates of the k th ray within the i th cluster.

Due to the high resolution in time and space at mmWave frequencies, the proposed channel model uses the clustering approach with each cluster consisting of several rays closely spaced in time and angular domains.

The stochastic properties of root-mean-square (RMS) delay spread (DS), AS, elevation spread (ES), as well as the inter-cluster and intracluster parameters will be calculated.

A. RMS DS

The RMS DS is calculated based on the estimated MPCs using the SAGE algorithm, only including values having MPC

power within a certain range from the peak value. Here, a 30-dB dynamic range is applied when computing the RMS DS. The RMS DS is then calculated as [38]

$$\tau_{rms} = \sqrt{\frac{\sum_i \sum_k |\alpha^{(i,k)}|^2 (\tau^{(i,k)})^2}{\sum_i \sum_k |\alpha^{(i,k)}|^2} - \left(\frac{\sum_i \sum_k |\alpha^{(i,k)}|^2 \tau^{(i,k)}}{\sum_i \sum_k |\alpha^{(i,k)}|^2} \right)^2}. \quad (8)$$

B. RMS AS

Both global AS (GAS) and cluster AS (CAS) are calculated using the SAGE estimates. The GAS is defined as the second-order moment of the azimuth power spectrum at the Tx side and is given by [38]

$$\phi_{rms} = \sqrt{\frac{\sum_i \sum_k |\alpha^{(i,k)}|^2 (\phi_{tx}^{(i,k)})^2}{\sum_i \sum_k |\alpha^{(i,k)}|^2} - \left(\frac{\sum_i \sum_k |\alpha^{(i,k)}|^2 \phi_{tx}^{(i,k)}}{\sum_i \sum_k |\alpha^{(i,k)}|^2} \right)^2}. \quad (9)$$

For the CAS, only those MPCs that contribute to the considered cluster are accounted. The CAS for i th cluster is given by [39]

$$\Phi_{rms}^{(i)} = \sqrt{\frac{\sum_k |\alpha^{(i,k)}|^2 (\phi_{tx}^{(i,k)})^2}{\sum_k |\alpha^{(i,k)}|^2} - \left(\frac{\sum_k |\alpha^{(i,k)}|^2 \phi_{tx}^{(i,k)}}{\sum_k |\alpha^{(i,k)}|^2} \right)^2}. \quad (10)$$

C. RMS ES

The global ES (GES) and cluster ES (CES) can be calculated using equations similar to (9) and (10), respectively.

Table III summarizes the Tx positions, scenarios, Tx–Rx distances, number of clusters, DS, GAS, GES, average CAS, and average CES introduced in the previous sections. In 12 Tx positions, there are six LOS scenarios and six NLOS scenarios. Here LOS refers to the scenario where there is an unobstructed path between the Tx and Rx antennas, whereas NLOS refers to the scenario where there are obstructions between Tx and Rx. This definition is in accordance with [40]. We can see from Table III that Tx1, Tx5, and Tx6 are LOS scenarios with boresight alignment, Tx2, Tx9, and Tx10 are within the Rx directional antenna's coverage, while at other Tx antenna positions, the LOS link is too weak and out of the Rx antenna's coverage. Tx–Rx distance ranges from 2 to 6 m. Most of the estimated 100 MPCs are within the 30-dB power range. The number of clusters for each location is between 4 and 7, while the rays in each cluster can vary from 4 to 40. The RMS DS for LOS scenarios is smaller than that of NLOS scenarios, which agrees with the previous studies [23]. Also, AS is much larger than ES. The average GAS and GES are 74.0° and 6.8°, respectively. The average CAS and CES considering all measurement results are 32.6° and 4.6°, respectively. An interdependence between the cluster parameters can be observed. When global angular spread is large, its cluster angular spread is usually large as well. But some exceptions exist such as at Tx1.

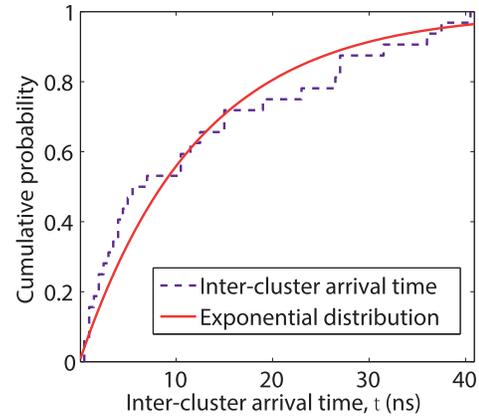


Fig. 18. CDF of the intercluster arrival times and an exponential distribution.

D. Intercluster Parameters

The intercluster arrival time is defined as the relative delay difference between two adjacent clusters. The delay of a cluster is taken as the delay of the strongest MPC in the cluster. Fig. 18 shows the cumulative distribution function (CDF) of the intercluster arrival rate. The distribution of the cluster arrival time is described by a Poisson process which is modeled by an exponential distribution [26]. The intercluster arrival rate Λ can be approximated by the inverse of the mean intercluster arrival time, which is 0.08 1/ns under both LOS and NLOS scenarios. It agrees with the result of 0.07 1/ns when using omnidirectional antenna at the Tx and 15° beamwidth antenna at the Rx in an office environment [26]. The cluster power decay rate Γ is defined as the decay rate of the strongest path within each cluster and modeled by an exponential decay. In our measurement, the power decay rate is estimated to be 2.5 ns under both LOS and NLOS scenarios.

E. Intracluster Parameters

The intracluster parameters of the channel model are estimated from the SAGE results. The intracluster characteristics in angular domain such as CAS and CES have been studied in previous sections. This section focuses on the intracluster characteristics in time domain. Based on the identified individual rays in time domain, statistical characteristics including average number of rays, ray arrival rate, and ray power decay time are calculated. The average number of rays in each cluster is about 20. But some clusters have the number of rays less than 20. The intracluster time-domain model was proposed in [25]. The cluster consists of a central ray $\alpha^{(i,0)}$ with fixed amplitude, precursor rays $\alpha^{(i,-N_{pre})}, \dots, \alpha^{(i,-1)}$, and postcursor rays $\alpha^{(i,1)}, \dots, \alpha^{(i,N_{post})}$. The central ray $\alpha^{(i,0)}$ is selected as the ray with highest amplitude in the cluster. The numbers of precursor rays N_{pre} and postcursor rays N_{post} can be derived from measurements. Precursor and postcursor rays are modeled as two Poisson processes with arrival rates λ_{pre} and λ_{post} , respectively. The average amplitudes A_{pre} and A_{post} of the precursor and postcursor rays decay exponentially with power decay times γ_{pre} and γ_{post} , respectively. They can

TABLE III
RESULT SUMMARY: CLUSTERS, DS, AND ANGULAR SPREAD

| Position | Scenario | Distance (m) | No. of cluster | DS (ns) | GAS (°) | GES (°) | Average CAS (°) | Average CES (°) |
|----------|----------|--------------|----------------|---------|---------|---------|-----------------|-----------------|
| Tx1 | LOS | 4.00 | 6 | 1.9 | 76.6 | 7.8 | 3.0 | 5.3 |
| Tx2 | LOS | 5.54 | 4 | 17.8 | 80.8 | 5.5 | 73.4 | 4.6 |
| Tx3 | NLOS | 4.24 | 7 | 19.8 | 66.7 | 6.4 | 35.1 | 4.5 |
| Tx4 | NLOS | 3.50 | 5 | 16.9 | 74.1 | 6.9 | 47.0 | 4.2 |
| Tx5 | LOS | 3.00 | 4 | 1.8 | 84.0 | 11.9 | 17.8 | 4.2 |
| Tx6 | LOS | 2.00 | 6 | 1.4 | 79.5 | 9.7 | 10.8 | 10.7 |
| Tx7 | NLOS | 3.08 | 5 | 9.7 | 62.5 | 7.0 | 49.1 | 4.1 |
| Tx8 | NLOS | 2.34 | 6 | 7.2 | 77.8 | 5.8 | 15.4 | 4.8 |
| Tx9 | LOS | 6.09 | 4 | 11.6 | 103.2 | 8.2 | 62.0 | 4.7 |
| Tx10 | LOS | 5.01 | 4 | 14.4 | 54.7 | 3.7 | 35.6 | 2.6 |
| Tx11 | NLOS | 4.00 | 6 | 15.2 | 55.8 | 4.4 | 26.7 | 2.0 |
| Tx12 | NLOS | 3.12 | 5 | 9.2 | 72.4 | 5.0 | 15.3 | 3.3 |
| Average | N/A | 3.83 | 5.1 | 10.6 | 74.0 | 6.8 | 32.6 | 4.6 |

TABLE IV
INTRA-CLUSTER TIME-DOMAIN PARAMETERS: COMPARISON WITH [23]
AND IEEE802.11AD CONFERENCE ROOM MODEL

| Parameter | Notation | LOS | NLOS | LOS/obstructed LOS [23] | 802.11ad |
|------------------|-------------------------|------|------|-------------------------|----------|
| Ray decay time | γ_{pre} [ns] | 4 | 3.9 | 4.6/4.8 | 3.7 |
| | γ_{post} [ns] | 5.4 | 4.5 | 4.7/4.5 | 4.5 |
| Ray K factor | K_{pre} [dB] | 11.5 | 3.3 | 8.6/10.3 | 10 |
| | K_{post} [dB] | 8.4 | 8.9 | 9.0/11.0 | 14.2 |
| Ray arrival rate | λ_{pre} [1/ns] | 0.5 | 0.7 | 0.9/1.1 | 0.37 |
| | λ_{post} [1/ns] | 0.5 | 1.2 | 0.9/1.0 | 0.31 |

be expressed as [25]

$$\begin{aligned} A_{pre}(\tau) &= A_{pre}(0)e^{|\tau|/\gamma_{pre}} \\ A_{post}(\tau) &= A_{post}(0)e^{-|\tau|/\gamma_{post}}. \end{aligned} \quad (11)$$

Rician K-factors are defined as the ratio of the amplitude of the central ray of the cluster $\alpha^{(i,0)}$ to the amplitudes of the precursor or postcursor rays

$$\begin{aligned} K_{pre} &= 20\log_{10} \left| \frac{\alpha^{(i,0)}}{A_{pre}(0)} \right| \text{ (dB)} \\ K_{post} &= 20\log_{10} \left| \frac{\alpha^{(i,0)}}{A_{post}(0)} \right| \text{ (dB)}. \end{aligned} \quad (12)$$

Table IV lists the estimated intracluster time domain parameters for the LOS and NLOS scenarios and compares them with the values from [23] and the IEEE802.11ad channel model [25] for a conference room environment. Although they are for a conference room, the size of our office room is close to theirs. Our results show larger values for the LOS postcursor ray decay time and smaller postcursor ray K factors, due to the differences in the measurement environment and the data processing method.

V. CONCLUSION

In this paper, a signal generator-VNA-based measurement system has been established in an office environment for 60-GHz indoor channels. Measurement results generated from the RDA and UVA methods have been compared and verified. It turns out that the RDA method highly relies on the accuracy

of antenna radiation pattern, while the UVA method is easier to implement and can get high precision and resolution in azimuth and elevation angular domains by using the large 3-D virtual antenna array. Measurement results with six LOS and six NLOS scenarios using the UVA method have been estimated using the SAGE algorithm, and clustered using the K-means algorithm.

The measured channel has been modeled based on an angular extended S-V model. Both intercluster and intracluster parameters have been characterized. Especially, angular characteristics including both azimuth and elevation domains in global level and cluster level have been investigated. We have found out that azimuth departure angles are diverse and highly related to antenna position and the measurement environment, while the elevation departure angles are more related to the antenna height difference and confined in a relatively small direction range. Meanwhile, the estimated AS is much larger than the ES in both global and cluster levels. Estimated parameters for cluster decay rate, cluster arrival rate, ray decay rate, Rician K-factor, and ray arrival rate have also been presented and compared with existing studies in a similar size conference room environment.

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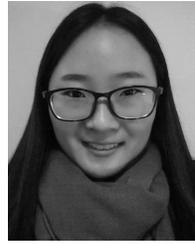
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